Gödel Spacetimes and *Flacuum* Solutions

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Seminar given on April 29th 2004 at University of Tel Aviv Based on hep-th/0401005. Work done in collaboration with *Patrick Meessen* (C.E.R.N.) Introduction/Motivation

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 - ▶ The BPST instanton configuration is realized in solutions with S^7 subspaces.

We are going to classify the maximally supersymmetric vacua of SUGRAs with 8 $Q_{\rm S}$ and find an interesting example of maximally supersymmetric, topologically non-trivial field configuration of SUGRA that corresponds to a well-known Abelian Yang-Mills instanton configuration.

Plan of the Talk:

- 1 SUGRA Vacua
- $6 \quad 8 \mathcal{Q} \text{ SUGRA Vacua}$
- 10 Timelike KK
- 13 The Flacuum
- 21 Conclusion



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- ★ Usually enjoys a high degree of (residual) symmetry. This symmetry determines all the kinematical properties of the QFT (conserved charges, spectrum etc.)
- ★ In (Special-Relativistic) QFT it is *required* that the residual symmetry of the vacuum includes the Poincaré group.

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Clearly, the most important question is

"How should (we or the theory) choose the vacuum?"

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This is a generalization of the concept of isometry, an infinitesimal general coordinate transformation generated by $\xi^{\mu}(x)$ that leaves the metric $g_{\mu\nu}$ invariant because it satisfies the *Killing (vector) equation*

$$\delta_{\xi}g_{\mu\nu} = 2\nabla_{(\mu}\xi_{\nu)} = 0.$$
(3)

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$$\xi^{\mu}_{(I)}(x) \to P_I \,,$$

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These will be the superalgebras of the QFTs constructed on these vacua!

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TheoryFieldsBosonic Action

d = 6, N = (1, 0)

d = 5, N = 1

d = 4, N = 2



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 d = 4, N = 2 $\{e^a{}_{\mu}, V_{\mu}, \psi_{\mu}\}$ $S = \int d^4 x \sqrt{|g|} \left[R - \frac{1}{4}F^2\right]$



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The maximally supersymmetric solutions of the three theories are related as follows:



- $aDS_3 \times S^3$ is the NHL of the extreme selfdual string. KG6 is the PL of $aDS_3 \times S^3$.
- $aDS_2 \times S^3$ is the NHL of the extreme black hole.
- $aDS_2 \star S^2$ is the NHL of the extreme rotating BMPV black hole.
- $aDS_3 \times S^2$ is the NHL of the extreme, critically rotating BMPV black hole and of the extreme string.
- $H_2 \star S^2$ is the NHL of the extreme overrotating BMPV black hole. (Fiol, Hofman, Lozano-Tellechea, hep-th/0312209)
- KG5 is the of the PL of the $aDS_n \times S^m$ families. G5 is the of a singular limit of the $H_2 \star S^2$ family. $aDS_2 \times S^2$ is the NHL of the extreme RN black hole. KG4 is the of the PL of $aDS_2 \times S^2$.

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 $aDS_2 \star S^2$ $ds^2 = -(d\psi + \omega)^2 + (R_3/2)^2 [d\Pi_{(2)}^2 - d\Omega_{(2)}^2],$ $F = -\frac{\sqrt{3}}{2} R_3 (\cos\xi \omega_a DS_2 - \sin\xi \omega_{(2)}).$ $\omega = R_3/2 (\cos\xi \cos\theta d\varphi + \sin\xi \cosh\chi dt).$

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$F = -\frac{\sqrt{3}}{2} R_3(\cos\xi\omega_{aDS_2} - \sin\xi\omega_{(2)}).$	$F = -\frac{\sqrt{3}}{2}R_3(\sinh\xi\omega_{H_2} + \cosh\xi\omega_{(2)}),$
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$\begin{array}{l} \left(\begin{array}{c} \text{G\"odel} \right) G \\ ds^2 \ = \ (dt + \omega)^2 - d\vec{x}_4^2 , \\ V \ = \ -\sqrt{3}\omega , \\ \omega \ = \ \lambda (x^1 dx^2 - x^3 dx^4) . \end{array}$	The spacelike fibrations over base spacetimes are used in standard KK reductions. ω becomes the $d = 4$ Maxwell field. Can we exploit timelike fibra- tions over a Euclidean space too?



It is possible to perform Kaluza-Klein dimensional reductions on timelike directions. The original (Lorentzian) theory is reduced to an Euclidean theory and its solutions (with timelike U(1) fibrations) are reduced to Euclidean solutions that may be interpreted as instantons.

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- We will deal only with Dirac fermions, but it is not always clear if we are dealing with vector or pseudovector fields, whose Wick rotations require an extra factor of i.



- There is no (known) Euclidean 8Q SUGRA in d = 6 (selfduality can't be Wick-rotated).
- There is only one way possible Wick rotation of the d = 5 theory if we want a real action.
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4 – The Flacuum

As we have seen, the dimensional reduction of the Gödel solution of d = 5, N = 1 SUGRA given by

$$\begin{array}{l} \left(\begin{array}{c} \mbox{G\"odel} \right) \ \mbox{G5} \end{array} \\ ds^2 \ = \ (dt + \omega)^2 - d\vec{x}_4^2 \, , \\ V \ = \ -\sqrt{3}\omega \, , \\ \omega \ = \ \lambda (x^1 dx^2 - x^3 dx^4) \, . \end{array} \end{array}$$

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leads to a non-trivial, maximally supersymmetric Euclidean solution of d = 4, N = 2 SUGRA (*i.e.* of the Einstein-Maxwell theory) with flat space and constant **anti-selfdual** field strength *F = -F ($F_{12} = -F_{34} = \lambda/2$)







A constant, anti-selfdual U(1) field strength certainly solves the Maxwell equation in flat space time, but,

how can flat space be a solution in presence of non-trivial matter?

The positivity properties of the action and the energy are opposite in Lorentzian and Euclidean signatures:

	Lorentzian	$\underline{\mathbf{Euclidean}}$
Action:	$-F^2 = \mathbf{E}^2 - \mathbf{B}^2$	$-F^2 = \mathbf{E}^2 + \mathbf{B}^2 > 0$
$T_{\mu u}$:	$F_{\mu}{}^{\rho}F_{\nu\rho} + {}^{\star}F_{\mu}{}^{\rho\star}F_{\nu\rho} > 0$	$F_{\mu}{}^{\rho}F_{\nu\rho} - {}^{\star}F_{\mu}{}^{\rho\star}F_{\nu\rho}$

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$$\underline{\text{Lorentzian}} \qquad \underline{\text{Euclidean}}$$

$$\underline{\text{Action:}} \qquad -F^2 = E^2 - B^2 \qquad -F^2 = E^2 + B^2 > 0$$

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In particular, selfdual and anti-selfdual Maxwell fields (that can only be defined in Euclidean signature) have a vanishing "energy-momentum" tensor. In general, (anti-) selfdual (non-) Abelian Yang-Mills configurations have vanishing energy-momentum tensors and <u>almost</u> decouple from the metric.

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The decoupling is not complete because (anti-) selfduality $F_{\rho\sigma} = \pm {}^{\star}F_{\rho\sigma}$ has to be proven w.r.t. to a given metric:

$$F_{\rho\sigma} = \pm \frac{1}{2\sqrt{|g|}} g_{\rho\mu} g_{\sigma\nu} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} \,.$$

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 $F = \pm^* F$ and $R_{\mu\nu} = \Lambda g_{\mu\nu}$, then $G_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{1}{2} T_{\mu\nu}$, and $\nabla_{\mu} F^{\mu\nu} = 0$

April 29th 2004

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The BPST SU(2) instanton

 $F = \pm^* F$ with any conformally flat metric. Since $F \to 0$ at ∞ we can take that of the round S^4

$$ds^2 = -\frac{d\vec{x}_4^2}{(1+(r/2R)^2)^2}, \quad \Rightarrow \quad R_{\mu\nu} = \frac{1}{R^2}g_{\mu\nu}.$$

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The flacuum U(1) solution

 $F = \pm^* F$ with any conformally flat metric. However, since F is constant, we have to stay with \mathbb{R}^4 which, at most, we can compactify on a torus to have a finite action. $R_{\mu\nu} = 0$ and the Einstein equation is satisfied with zero cosmological constant. Observe that taking the gauge group as U(1) is equivalent to take the time periodic in the Gödel solution.

The Hopf bundle $S^5 \xrightarrow{S^1} \mathbb{CP}^2$

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 S^{2n+1} is the complex hypersurface in \mathbb{C}^{n+1} with equation $z^{\alpha} \bar{z}^{\alpha} = 1$, $\alpha = 1, \dots, n, \sharp$.

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 $\implies ds^2_{\mathbb{CP}^n} = g_{i\bar{j}}d\xi^i \otimes d\bar{\xi}^j + g_{\bar{i}j}d\bar{\xi}^i \otimes d\xi^j$ is the Hermitean Fubini-Study metric on \mathbb{CP}^n .

The Hopf bundle $S^5 \xrightarrow{S^1} \mathbb{CP}^2$

 S^{2n+1} is the complex hypersurface in \mathbb{C}^{n+1} with equation $z^{\alpha} \bar{z}^{\alpha} = 1$, $\alpha = 1, \dots, n, \sharp$. It is convenient to change coordinates $z^{\alpha} \to u, \rho, \xi^{i}, i = 1, \dots, n$

$$\rho = |z^{\sharp}|, \quad u = z^{\sharp}/\rho, \quad \xi^{i} = z^{i}/z^{\sharp}, \quad (\leftarrow \text{ projective coordinates in } \mathbb{CP}^{n})$$

in which S^{2n+1} is $\rho = \frac{1}{1 + \xi^{i} \overline{\xi^{i}}}.$

Now we substitute the S^{2n+1} equation into the Euclidean metric on \mathbb{C}^{n+1} , $ds^2 = dz^{\alpha} d\bar{z}^{\alpha}$, to find the metric of the round S^{2n+1} .

It takes the form

$$ds^2_{\mathbf{S}^{2n+1}} = ds^2_{\mathbb{CP}^n} + \omega^2 \,,$$

⇒ $ds_{\mathbb{CP}^n}^2 = g_{i\bar{j}}d\xi^i \otimes d\bar{\xi}^j + g_{\bar{i}j}d\bar{\xi}^i \otimes d\xi^j$ is the Hermitean Fubini-Study metric on \mathbb{CP}^n . ⇒ $\omega = u^{-1}du + A$ where A is a U(1) connection on \mathbb{CP}^n such that

$$dA = ig_{i\bar{\jmath}}d\xi^i \wedge d\bar{\xi}^j \equiv K \,,$$

the Kähler 2-form K, which is, therefore, closed $dK = d^2A = 0$.

»→ K, is also co-closed ${}^{\star}K = 0$, so \mathbb{CP}^n is Kähler and K therefore solves the Maxwell equations on \mathbb{CP}^n (Trautman, 1977).

End of slide

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- ▶ The components T_{ij} and $T_{\bar{i}\bar{j}}$ trivially vanish:

$$T_{ij} = K_{i\bar{k}}K_{j\bar{l}}g^{\bar{k}\bar{l}} - g_{ij}K^2 = 0,$$

and the components $T_{i\bar{j}}$ and $T_{\bar{i}j}$ vanish for n = 2:

$$T_{i\bar{\jmath}} = K_{i\bar{k}} K_{l\bar{\jmath}} g^{\bar{k}l} - \frac{1}{4} g_{i\bar{\jmath}} (2K_{k\bar{l}} g^{\bar{l}m} K_{m\bar{n}} g^{\bar{n}k}) = -g_{i\bar{\jmath}} + \frac{1}{4} g_{i\bar{\jmath}} 2n \,.$$

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Then, since the Fubini-Study metric solves the Einstein equations with cosmological constant $\Lambda = +6$, we have another solution of the Euclidean Einstein-Maxwell equations. The embedding of this solution and the BPST instanton in supregravity are problematic.

Other solutions with vanishing Euclidean energy-momentum tensor can be obtained by time-like compactification of other Gödel solutions.

The vector field of our solution (in a new gauge)

$$V = \lambda (x^{1} dx^{2} - x^{2} dx^{1} - x^{3} dx^{4} + x^{4} dx^{3}) \equiv F_{ab} x^{a} dx^{b},$$

is not strictly periodic on T^4 : when we move around the *a*-th period from x to $x + \hat{a}$ it changes by a gauge transformation

$$V(x+\hat{a}) = V(x) + d\Lambda_a(x), \qquad \Lambda_a(x) = l^{(a)} F_{(a)b} x^b,$$

where $\Lambda_a(x)$ are the U(1) parameters, defined modulo 2π .

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Consistency requires that $V(x + \hat{a} + \hat{b}) = V(x + \hat{b} + \hat{a})$, that is

$$\Lambda_a(x+\hat{b}) + \Lambda_b(x) = \Lambda_b(x+\hat{a}) + \Lambda_a(x) \mod(2\pi),$$

which in our case implies

$$\lambda l^1 l^2 = \pi n, \qquad \lambda l^3 l^4 = \pi m,$$

for two integers n, m that label the possible non-trivial bundles.

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for two integers n, m that label the possible non-trivial bundles. The Euclidean action of the SUGRA solutions is

$$S = -4\pi^2 |\mathbf{nm}| \,.$$

Consistency requires that the Killing spinor can be identified with itself after a translation around one of the periods:

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Its has been argued that (Duff, Lu, Hull, Papadopoulos, Tsimpis) whant should be considered is the generalized holonomy of the supergravity theory, which is basically that of the gravitino supersymmetry transformation rule (the Killing spinor equation).

In this sense, the above transformations belong to the generalized holonomy group of N = 2, d = 4 SUGRA which is $SL(2, \mathbb{H})$ (Batrachenko, Wen hep-th/0402141).

The symmetry superalgebra of the flacuum solution is particularly interesting because it is a deformation of the supertranslation algebra that preserves the commutativity of momenta but modifies slightly the anticommutator of the supercharges (Berkovits and Seiberg)

$$\left\{ \begin{array}{lll} \mathcal{Q}_{(\alpha)}^{\dagger}, \mathcal{Q}_{(\beta)} \right\} &= (\gamma^{1} \gamma^{a})_{\alpha \beta} P_{(a)} & - [\gamma^{1} \frac{1}{2} (1 - \gamma_{5})]_{\alpha \beta} M , \\ \\ \left[\mathcal{Q}_{(\alpha)}, P_{(a)} \right] &= -\mathcal{Q}_{(\beta)} \Gamma_{s} (P_{(a)})^{\beta}{}_{\alpha} , \\ \\ \left[\mathcal{Q}_{(\alpha)}, M \right] &= -\mathcal{Q}_{(\beta)} \Gamma_{s} (M)^{\beta}{}_{\alpha} , \\ \\ \left[P_{(a)}, M \right] &= -P_{(b)} \Gamma_{v} (M)^{b}{}_{a} , \end{array}$$

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$$\begin{split} \left\{ \mathcal{Q}_{(\alpha)}^{\dagger}, \mathcal{Q}_{(\beta)} \right\} &= (\gamma^{1} \gamma^{a})_{\alpha \beta} P_{(a)} + (\gamma^{1} \gamma_{5})_{\alpha \beta} P_{(0)} - [\gamma^{1} \frac{1}{2} (1 - \gamma_{5})]_{\alpha \beta} M, \\ \left[\mathcal{Q}_{(\alpha)}, P_{(a)} \right] &= -\mathcal{Q}_{(\beta)} \Gamma_{s} (P_{(a)})^{\beta} \alpha, \\ \left[\mathcal{Q}_{(\alpha)}, M \right] &= -\mathcal{Q}_{(\beta)} \Gamma_{s} (M)^{\beta} \alpha, \\ \left[P_{(a)}, M \right] &= -P_{(b)} \Gamma_{v} (M)^{b} \alpha, \qquad \left[P_{(a)}, P_{(b)} \right] = F_{ab} P_{(0)}. \end{split}$$

The quantization of the string on this background leads to a non-commutative Field Theory in which only the fermionic superspace coordinates anticommute anomalously.

This superalgebra can be obtained by dimensional reduction of the Gödel superalgebra, in which the momenta $P_{(a)}$ do not commute, but give $P_{(0)}$ which should be interpreted as the generator of U(1) gauge transformations on d = 4. This property is, precisely, what allowed us to relate the periods of the torii.





5 – Conclusion

 \star We completed the classification of Lorentzian and Euclidean maximally supersymmetric vacua with 8 supercharges.





★ We have found a solution, the *flacuum* solution with very interesting properties and that can be generalized to other dimensions (always as a timelike reduction of a Gödel-type solution).

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- ★ We have discussed how the compactification affects the residual supersymmetry of the solution, which is a delicate point because the holonomy of the solution is not contained in SO(4).
- ★ We have determined the symmetry superalgebra of the *flacuum* solution. We notice that the symmetry superalgebras of all the maximally supersymmetric vacua are always deformations of the supertranslation (superPoincaré) algebra, which may allow to classify and find all these vacua.

