Gödel Spacetimes and *Flacuum* Solutions

Tomás Ortín (I.F.T.-C.S.I.C)

Seminar given on February 17th 2004 at IFT-UAM/CSIC Based on hep-th/0401005. Work done in collaboration with *Patrick Meessen* (C.E.R.N.) Introduction/Motivation

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 - ▶ The BPST instanton configuration is realized in solutions with S^7 subspaces.

In this seminar we are going to study an interesting example of maximally supersymmetric, topologically non-trivial field configuration of supergravity that corresponds to a well-known Abelian Yang-Mills instanton configuration.

Plan of the Talk:

- 1 SUGRA Vacua
- $6 \quad 8\mathcal{Q} \text{ SUGRA Vacua}$
- 10 Timelike KK
- 13 The Flacuum
- 18 Conclusion



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Gödel Spacetimes and Flacuum Solutions





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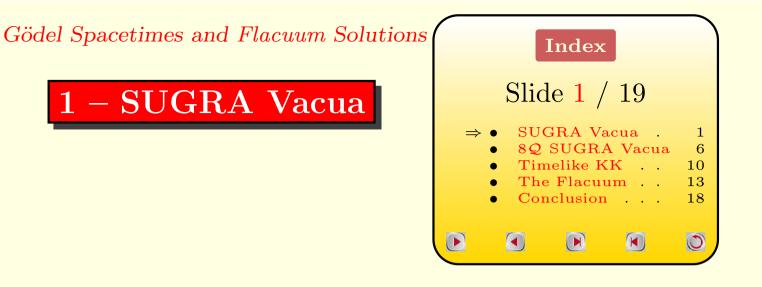




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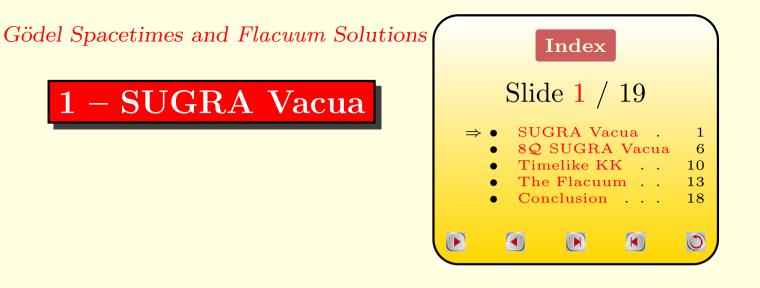
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- ★ Usually enjoys a high degree of (residual) symmetry. This symmetry determines all the kinematical properties of the QFT (conserved charges, spectrum etc.)
- ★ In (Special-Relativistic) QFT it is *required* that the residual symmetry of the vacuum includes the Poincaré group.

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Clearly, the most important question is

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This is a generalization of the concept of isometry, an infinitesimal general coordinate transformation generated by $\xi^{\mu}(x)$ that leaves the metric $g_{\mu\nu}$ invariant because it satisfies the *Killing (vector) equation*

$$\delta_{\xi} g_{\mu\nu} = 2\nabla_{(\mu} \xi_{\nu)} = 0.$$
 (3)

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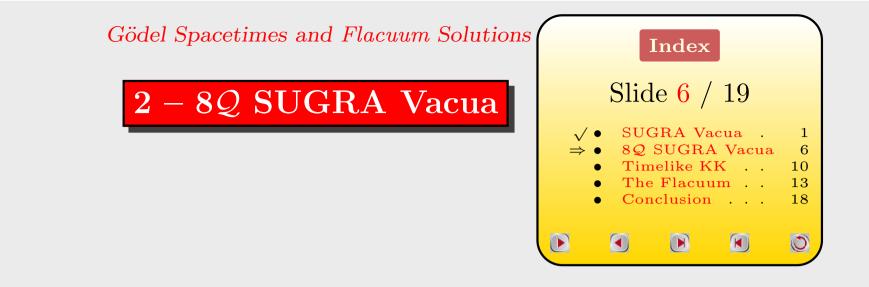
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TheoryFieldsBosonic Action

d = 6, N = (1, 0)

d = 5, N = 1

d = 4, N = 2

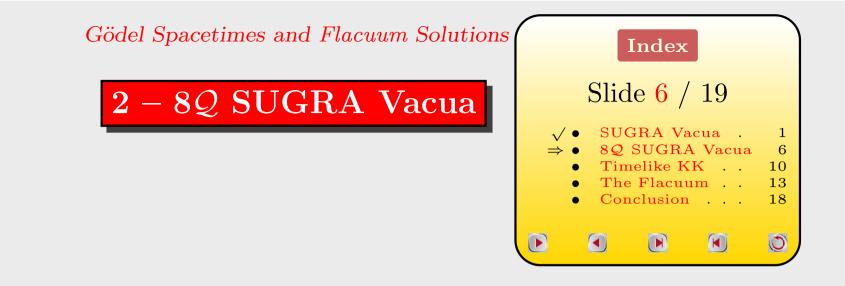


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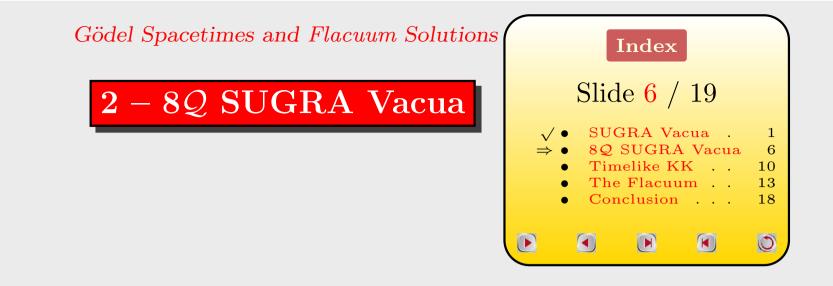
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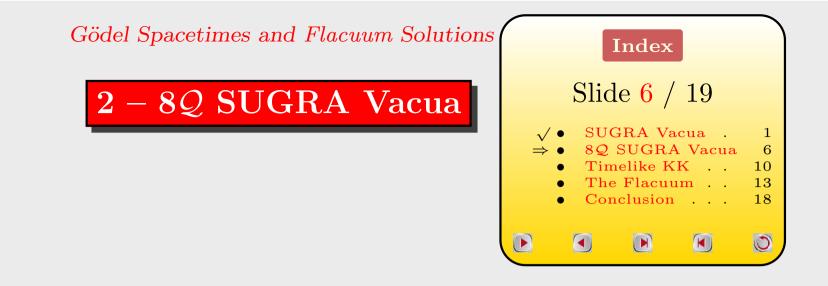
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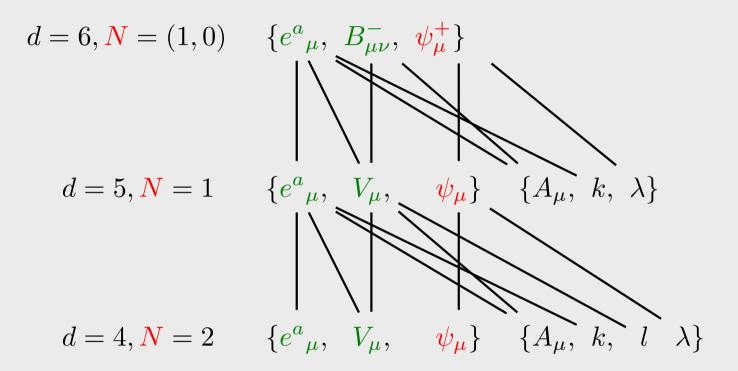
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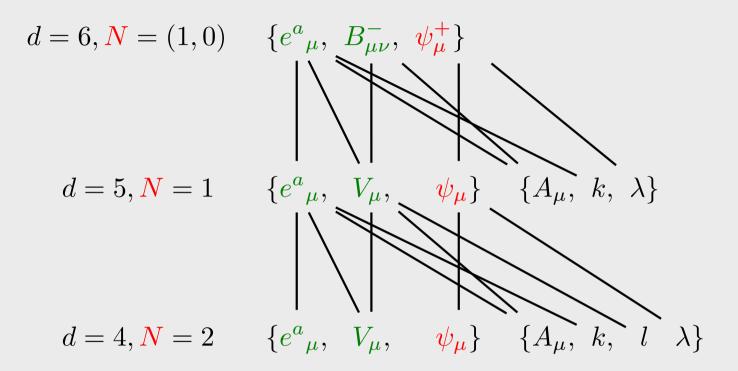
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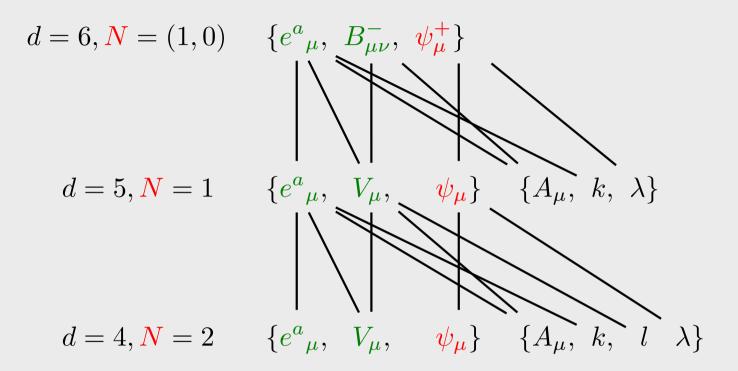
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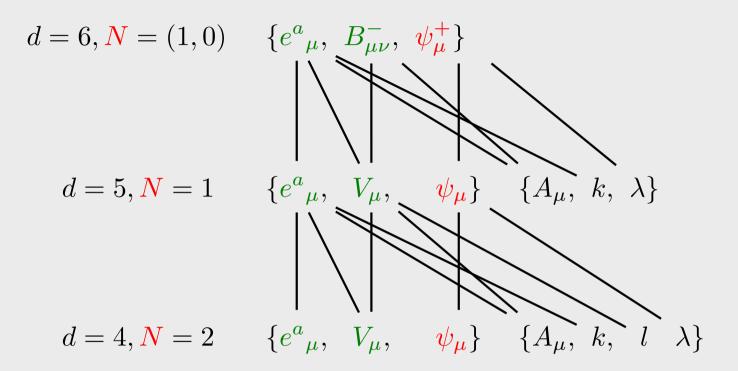
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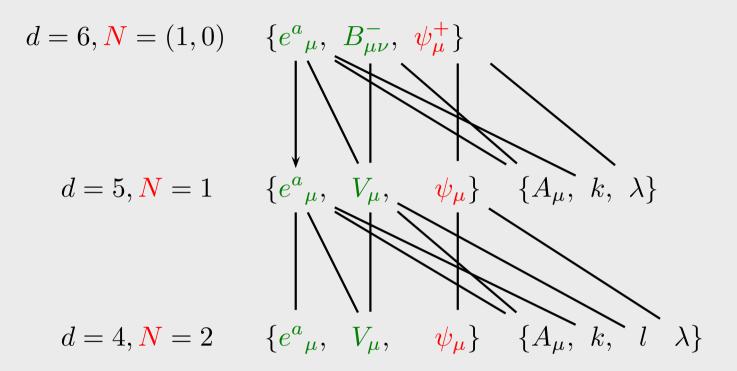
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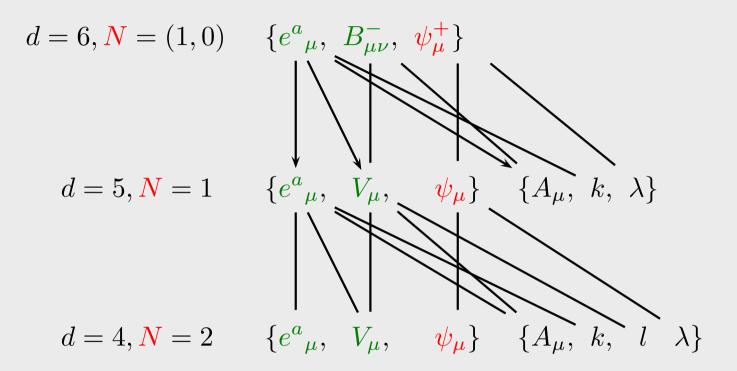
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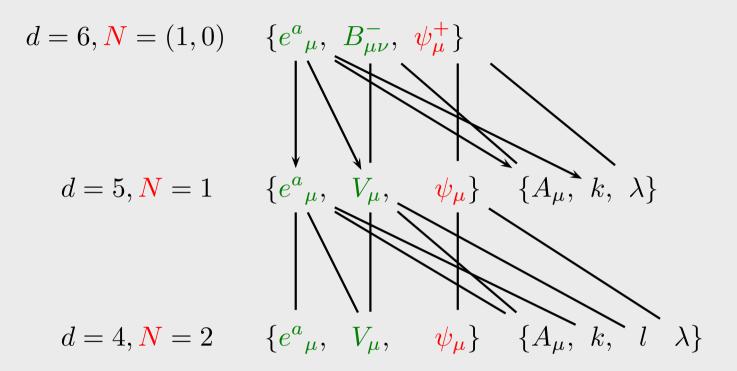
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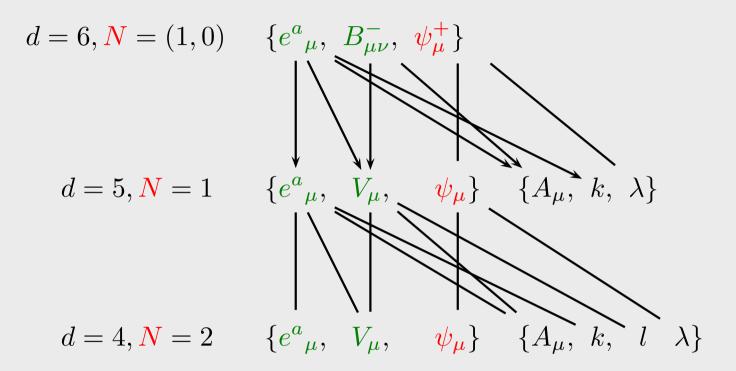
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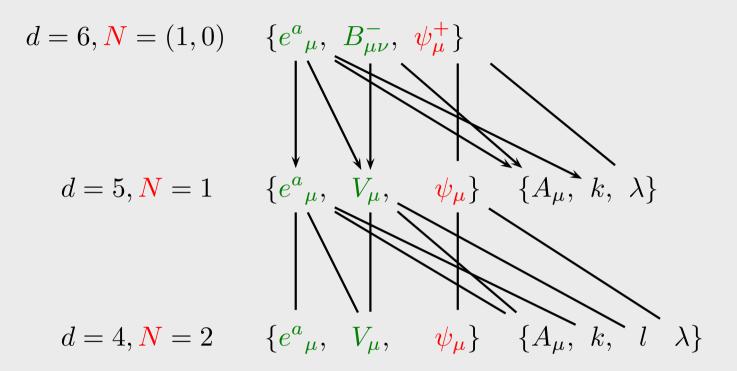
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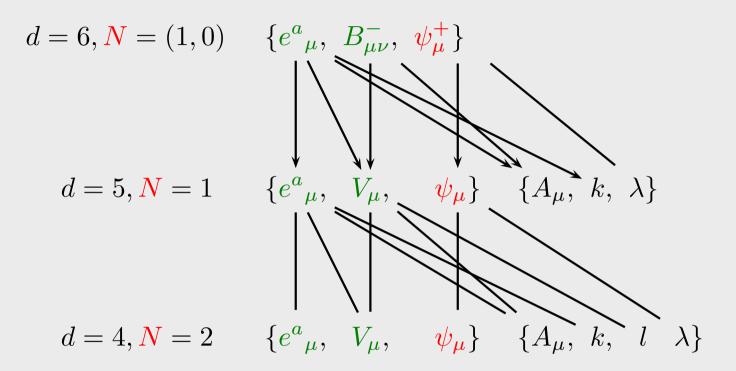
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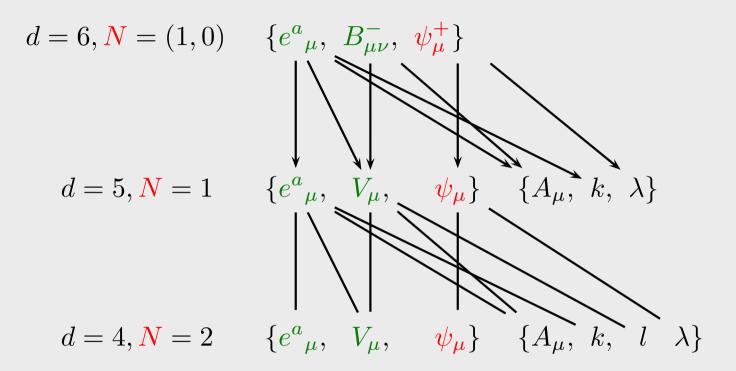
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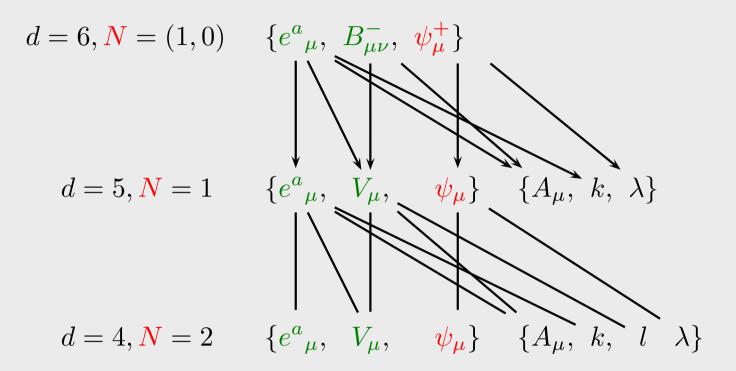


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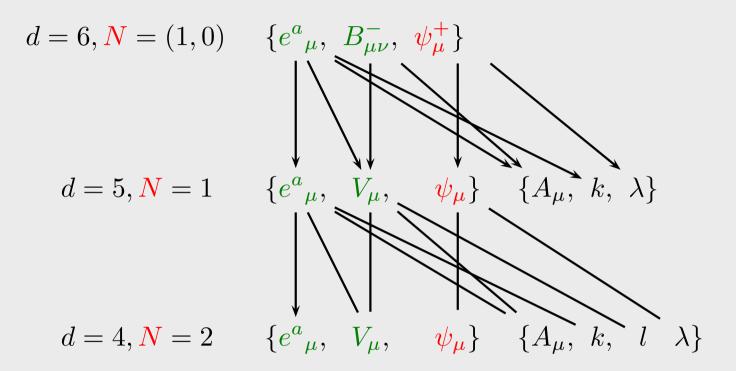


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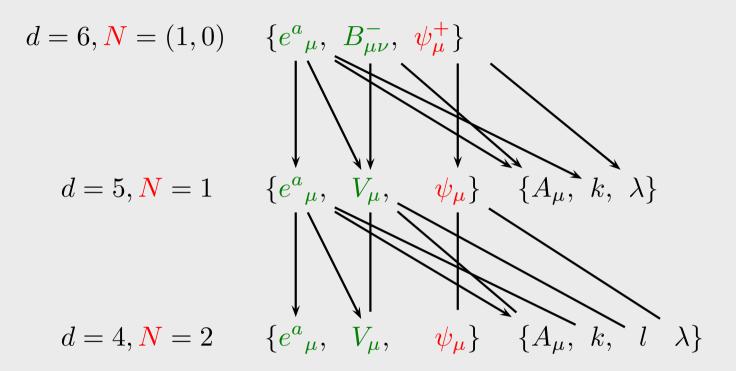
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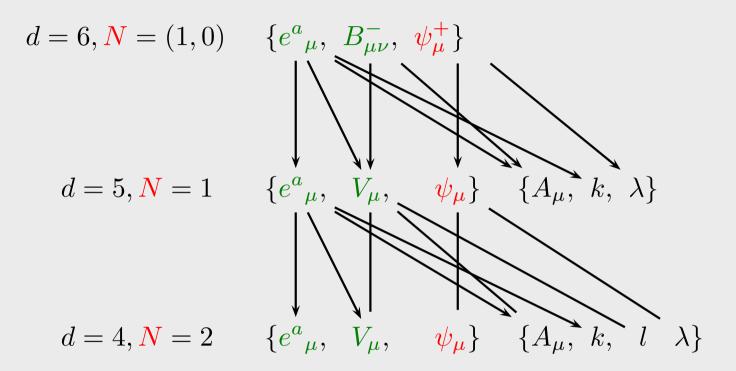
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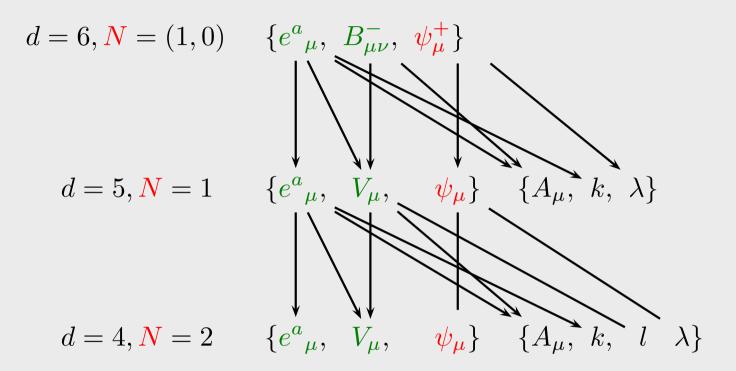
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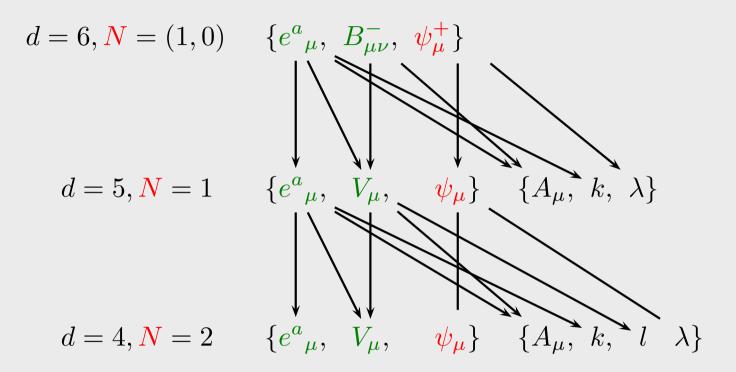
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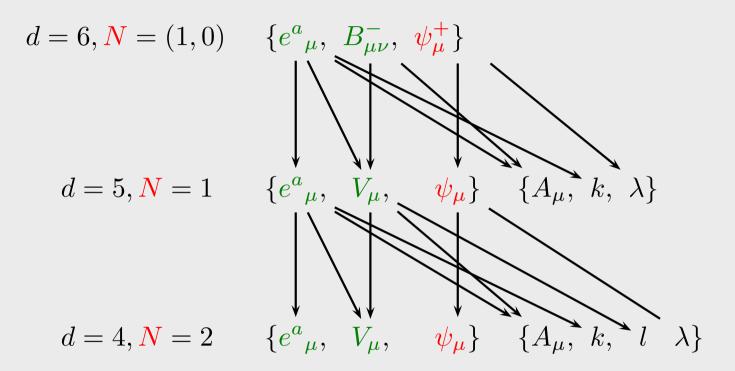


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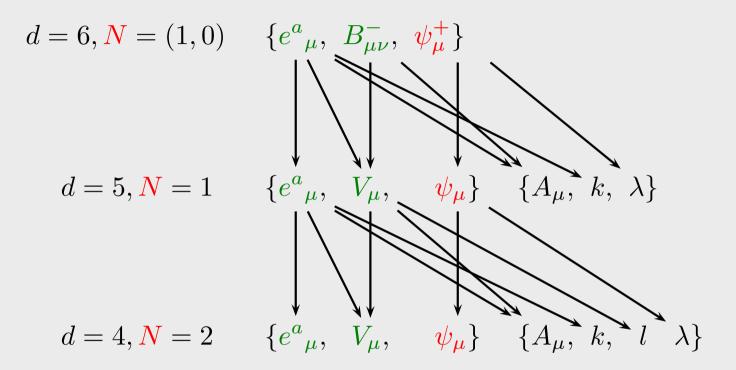


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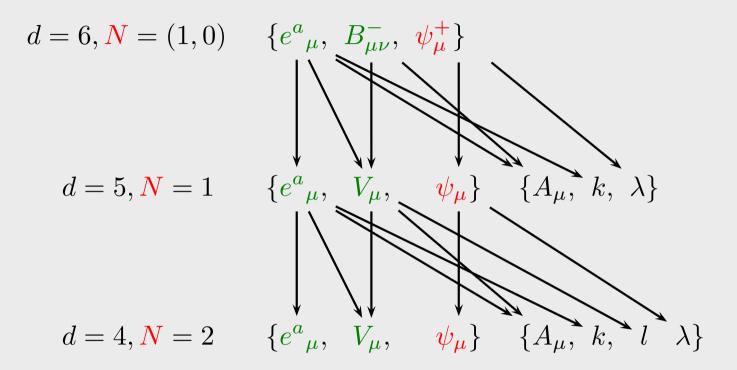
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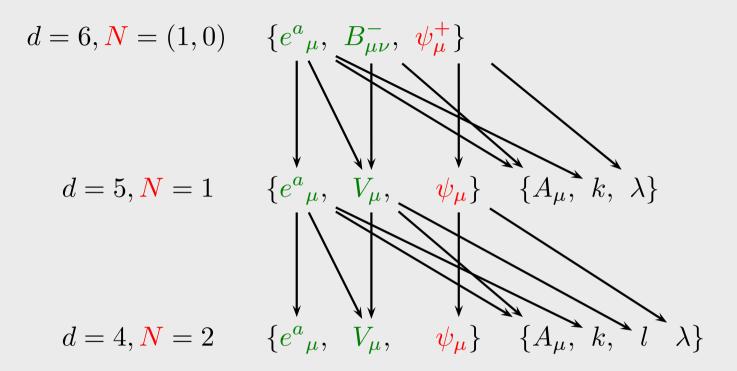
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February 20th 2004

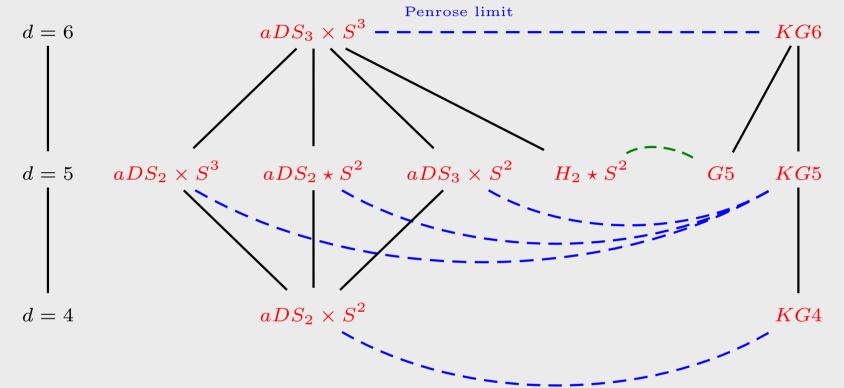


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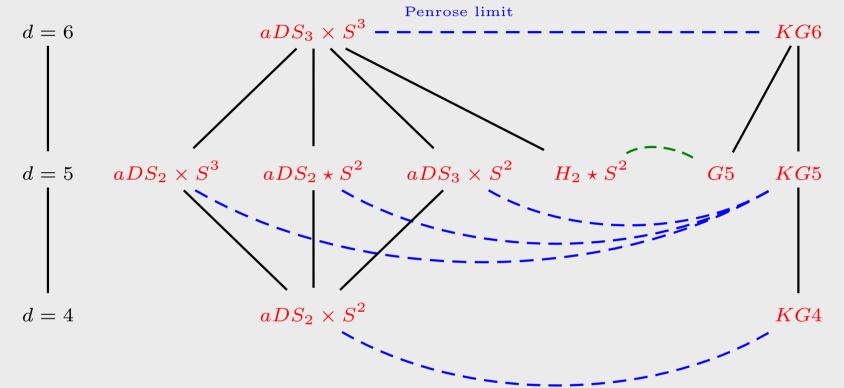
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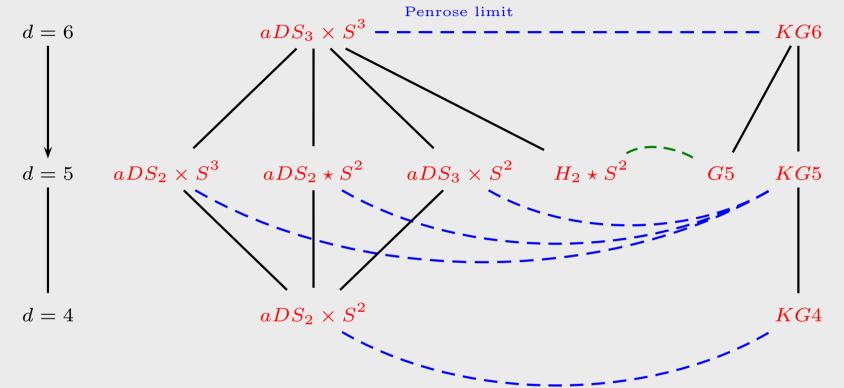


- $aDS_3 \times S^3$ is the NHL of the extreme selfdual string. KG6 is the PL of $aDS_3 \times S^3$.
- $aDS_2 \times S^3$ is the NHL of the extreme black hole.
- $aDS_2 \star S^2$ is the NHL of the extreme rotating BMPV black hole.
- $aDS_3 \times S^2$ is the NHL of the extreme, critically rotating BMPV black hole and of the extreme string.
- $H_2 \star S^2$ is the NHL of the extreme overrotating BMPV black hole.
- KG5 is the of the PL of the $aDS_n \times S^m$ families. G5 is the of a singular limit of the $H_2 \star S^2$ family. $aDS_2 \times S^2$ is the NHL of the extreme RN black hole. KG4 is the of the PL of $aDS_2 \times S^2$.

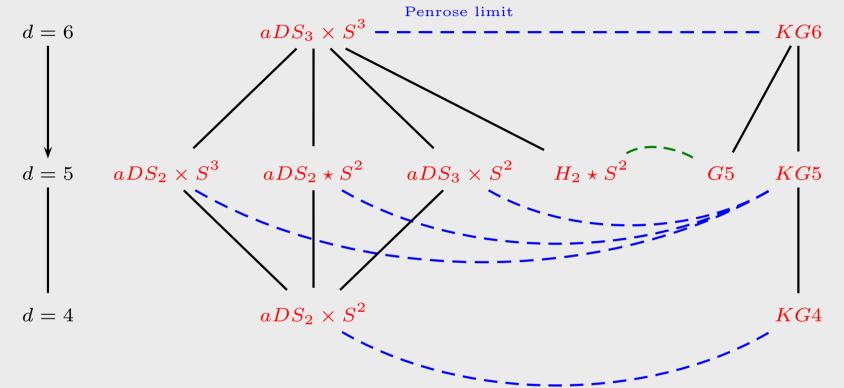
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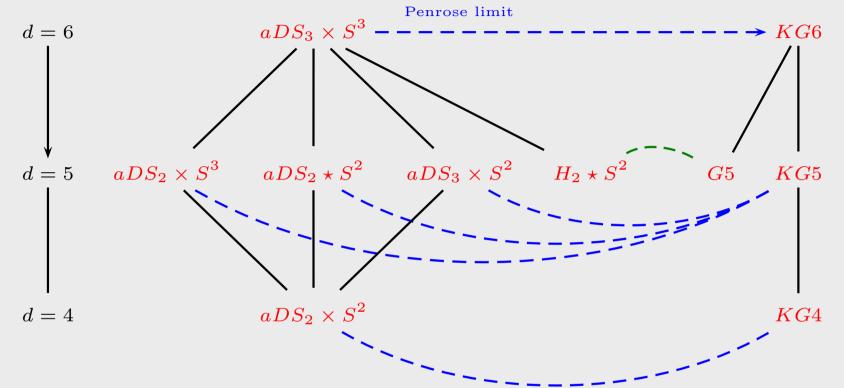
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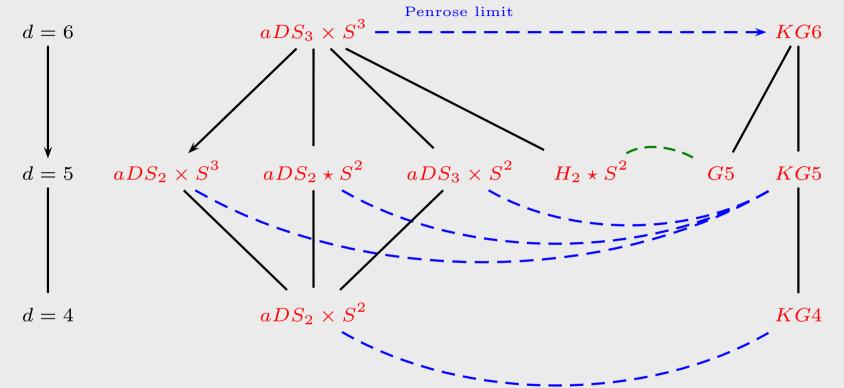
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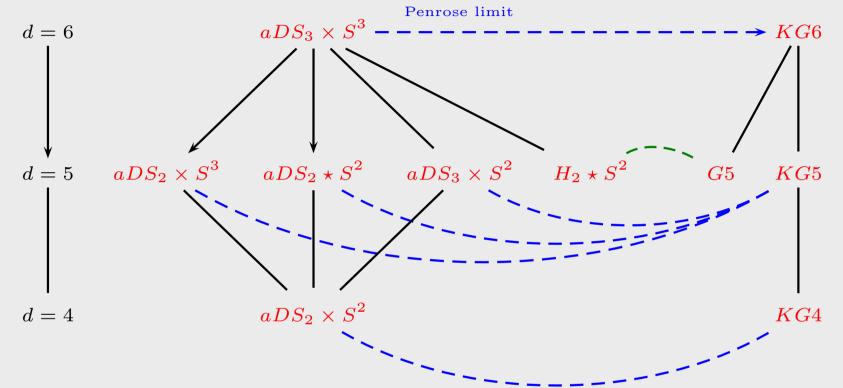
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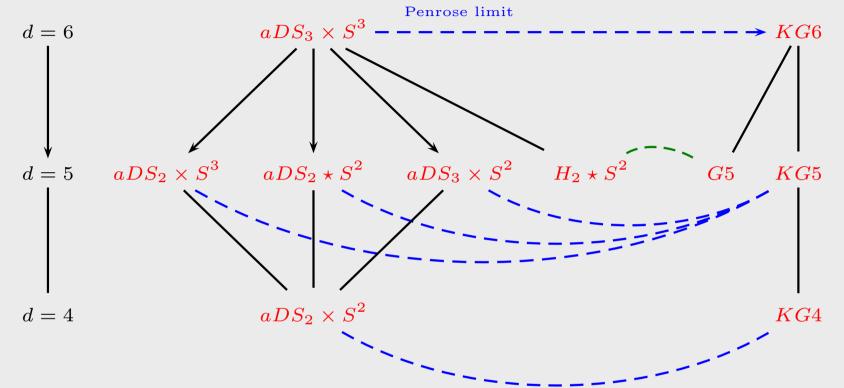
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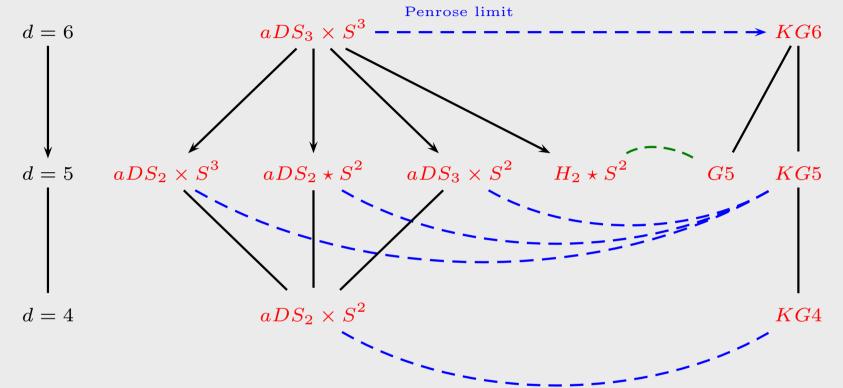


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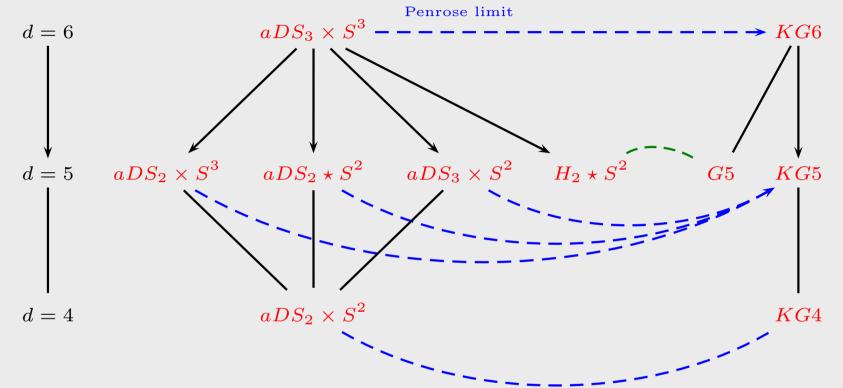
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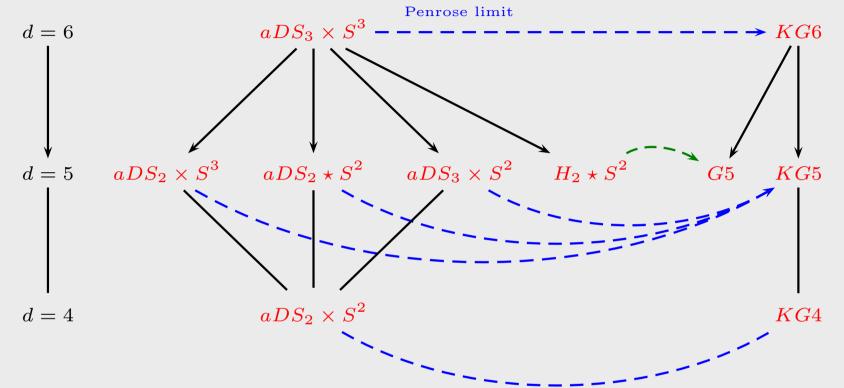


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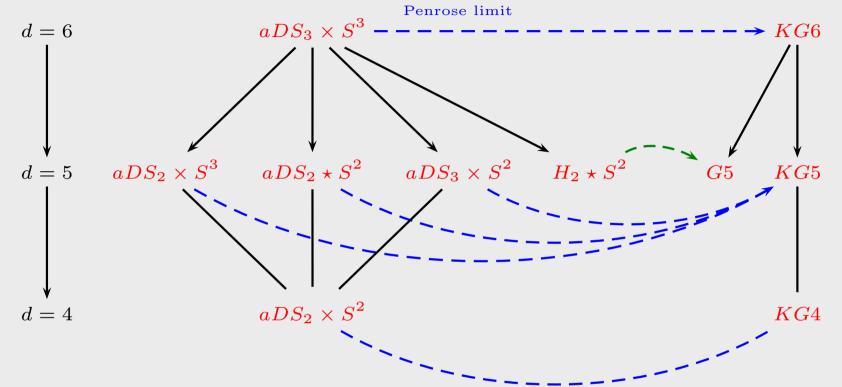
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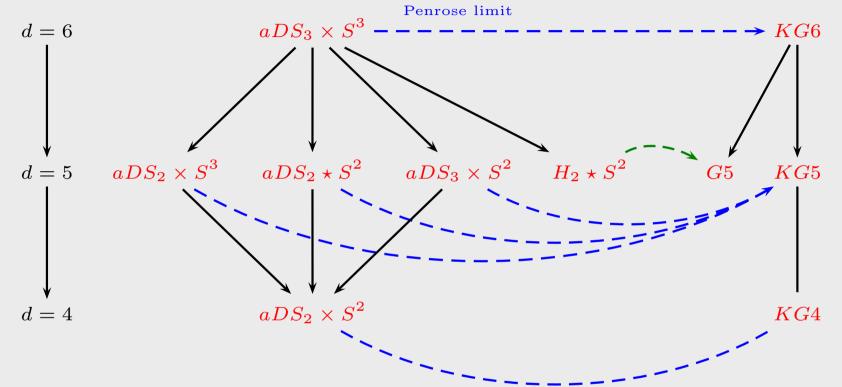


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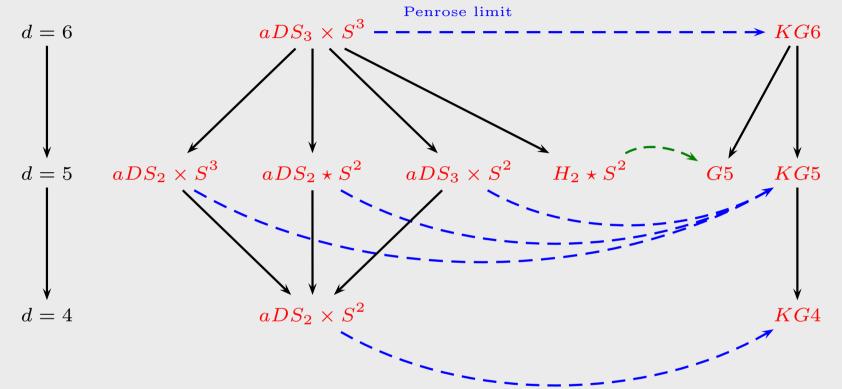


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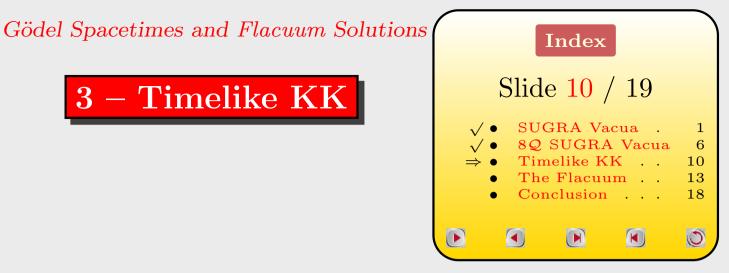


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 $aDS_2 \star S^2$ $ds^2 = -(d\psi + \omega)^2 + (R_3/2)^2 [d\Pi_{(2)}^2 - d\Omega_{(2)}^2],$ $F = -\frac{\sqrt{3}}{2} R_3 (\cos\xi \omega_{aDS_2} - \sin\xi \omega_{(2)}).$ $\omega = R_3/2 (\cos\xi \cos\theta d\varphi + \sin\xi \cosh\chi dt).$

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$F = -\frac{\sqrt{3}}{2} R_3 (\cos \xi \omega_{aDS_2} - \sin \xi \omega_{(2)}).$	$F = -\frac{\sqrt{3}}{2}R_3(\sinh\xi\omega_{H_2} + \cosh\xi\omega_{(2)}),$
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$\begin{array}{l} \left(\begin{array}{c} \mbox{G\"odel} \right) \ \mbox{G5} \end{array} \\ ds^2 \ = \ (dt + \omega)^2 - d\vec{x}_4^2 , \\ V \ = \ -\sqrt{3}\omega , \\ \omega \ = \ \lambda (x^1 dx^2 - x^3 dx^4) . \end{array} \end{array}$	The spacelike fibrations over base spacetimes are used in standard KK reductions. ω becomes the $d = 4$ Maxwell field. Can we exploit timelike fibra- tions over a Euclidean space too?



It is possible to perform Kaluza-Klein dimensional reductions on timelike directions. The original (Lorentzian) theory is reduced to an Euclidean theory and its solutions (with timelike U(1) fibrations) are reduced to Euclidean solutions that may be interpreted as instantons.

Gödel Spacetimes and Flacuum Solutions Index Slide 10 / 19 3 – Timelike KK SUGRA Vacua . 1 80 SUGRA Vacua 6 Timelike KK 10The Flacuum 1318Conclusion . . . 0

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Gödel Spacetimes and Flacuum Solutions





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Gödel Spacetimes and Flacuum Solutions

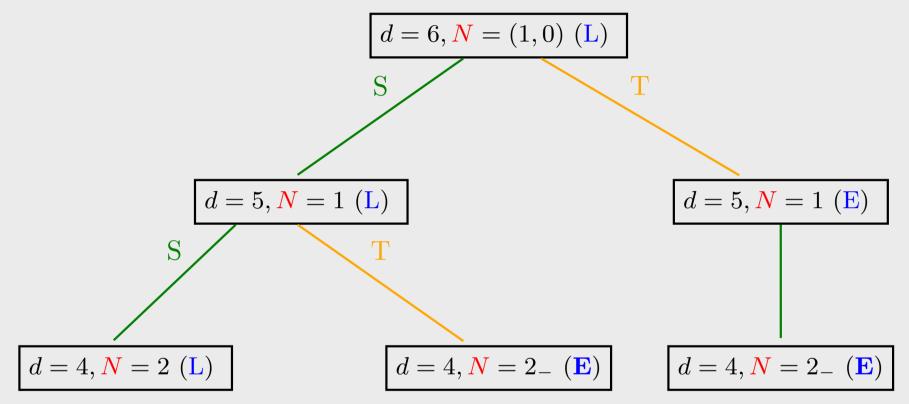




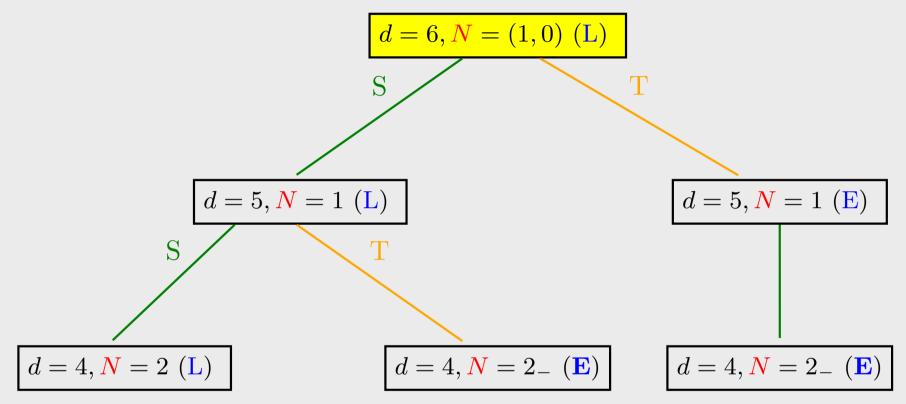
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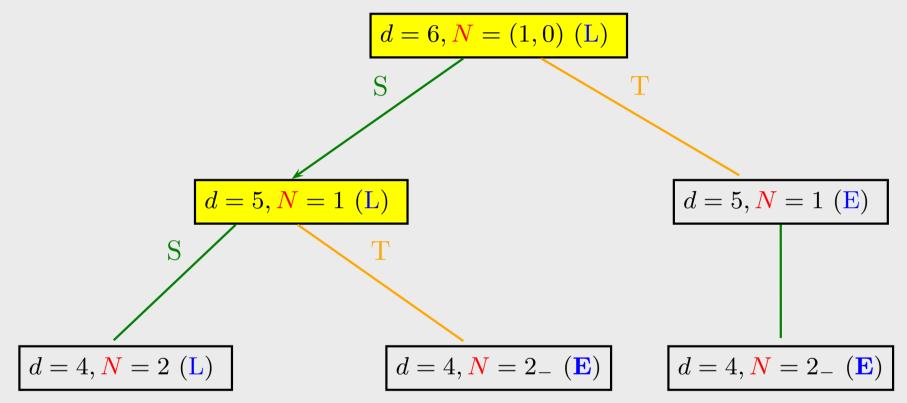
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- We will deal only with Dirac fermions, but it is not always clear if we are dealing with vector or pseudovector fields, whose Wick rotations require an extra factor of i.



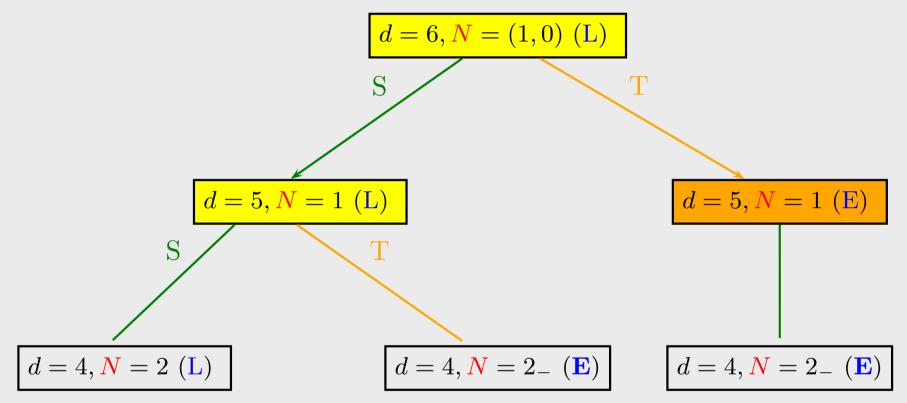
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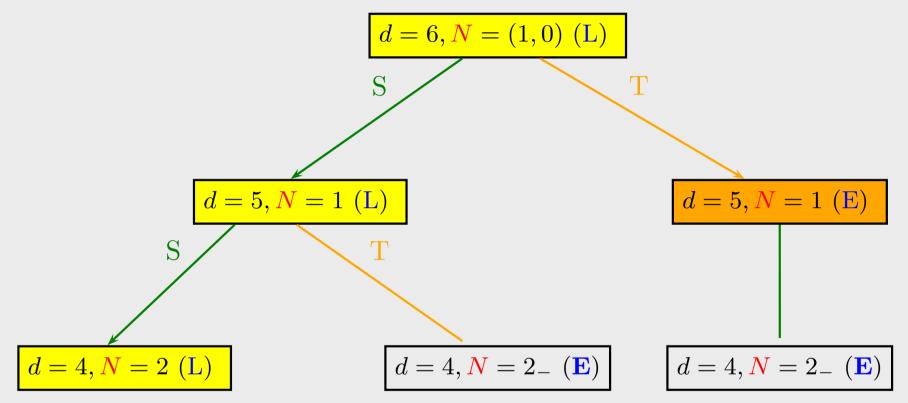
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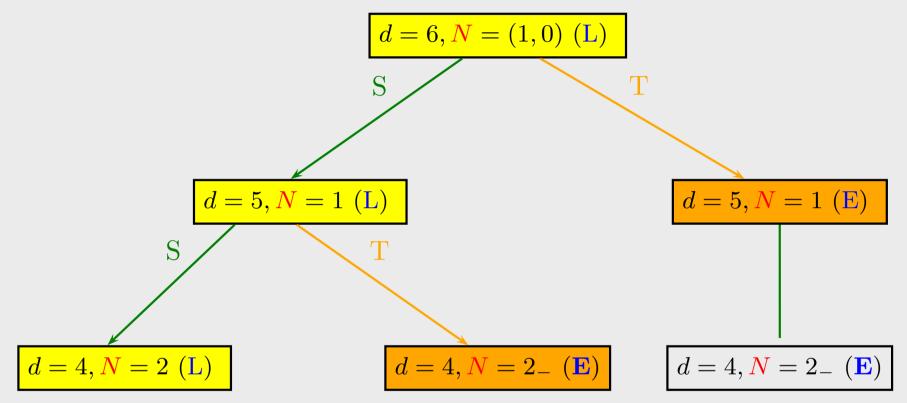
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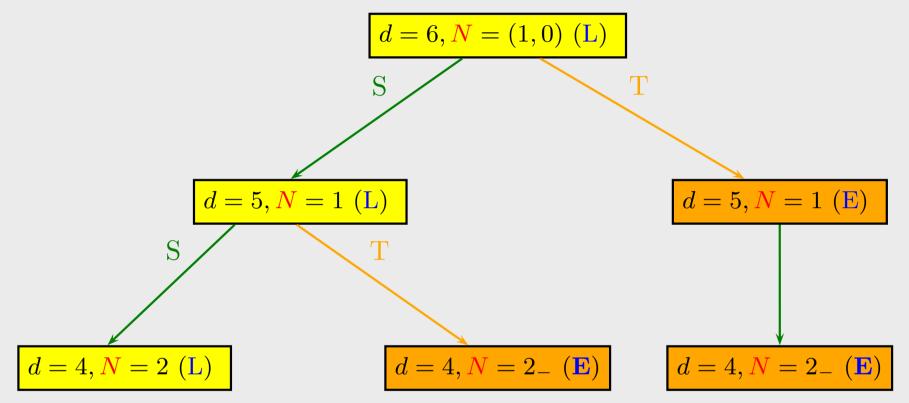
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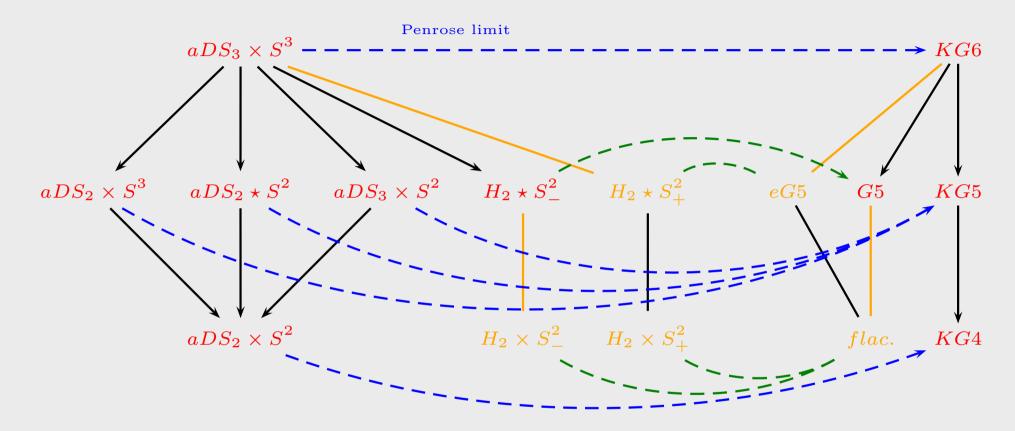
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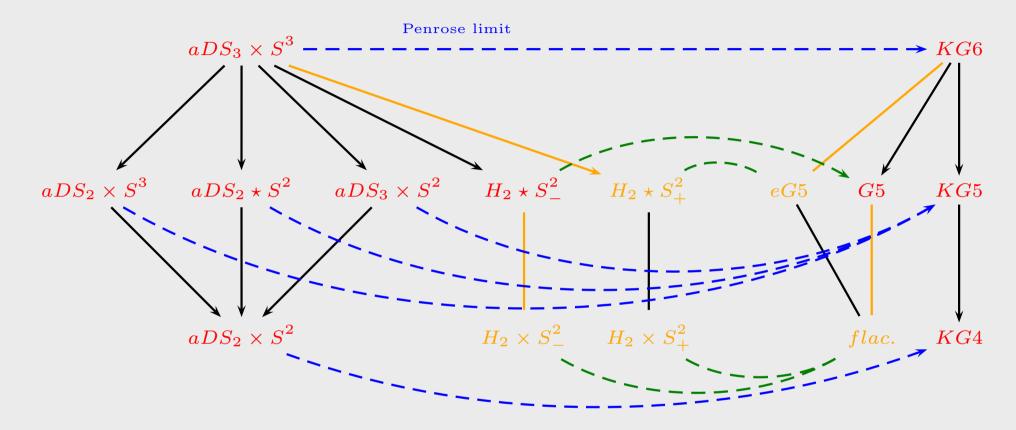


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eG5 is a Euclidean version of the Gödel spacetime G5. It can also be obtained by a singular limit procedure from $H_2 \star S_+^2$.

 $H_2 \times S_-^2 H_2 \times S_+^2$ are solutions of different theories and are related by analytical continuation.

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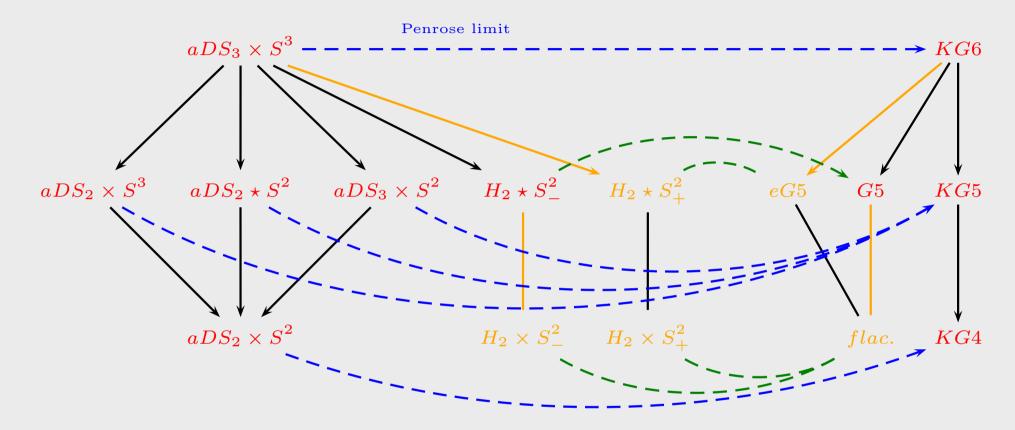


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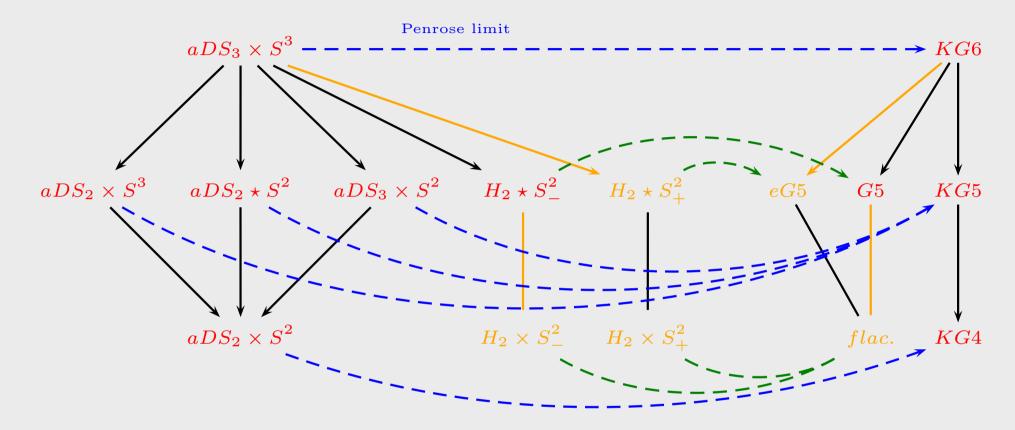


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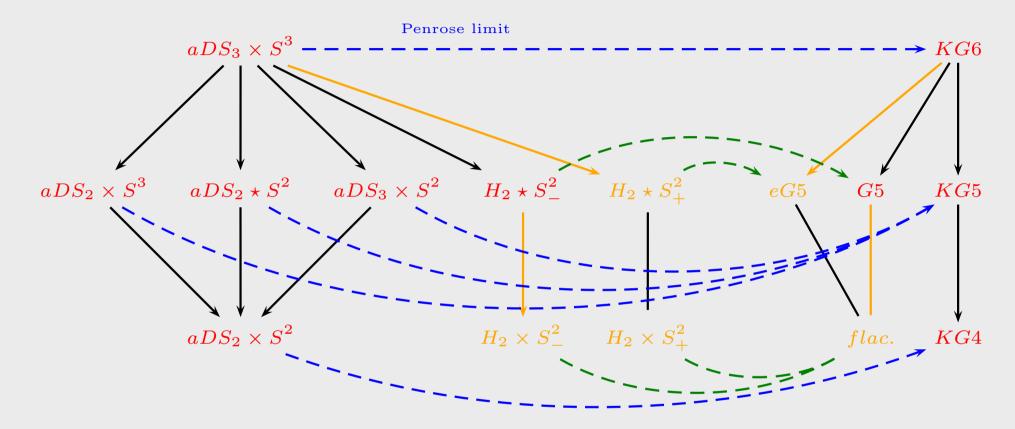


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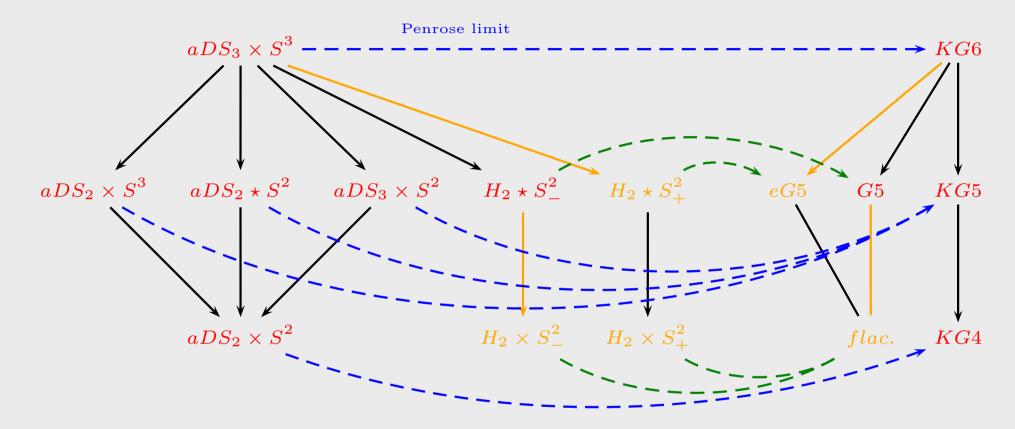


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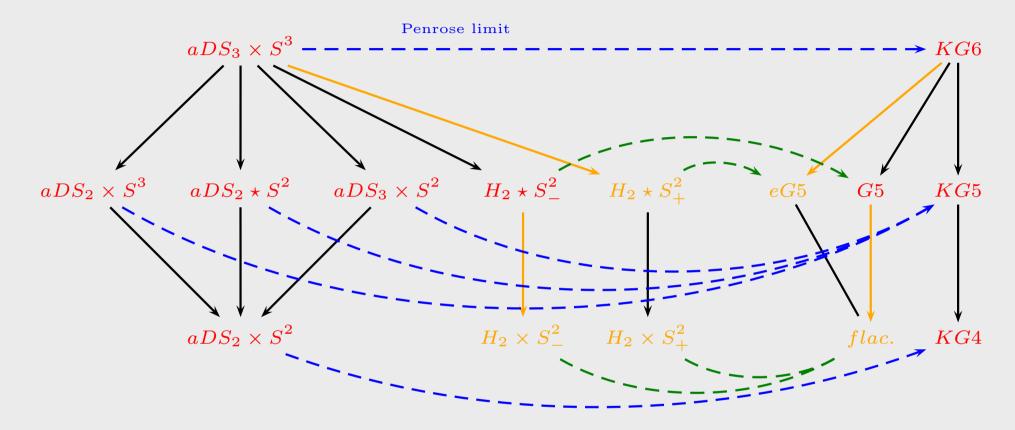
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Now we get new Euclidean solutions as well:



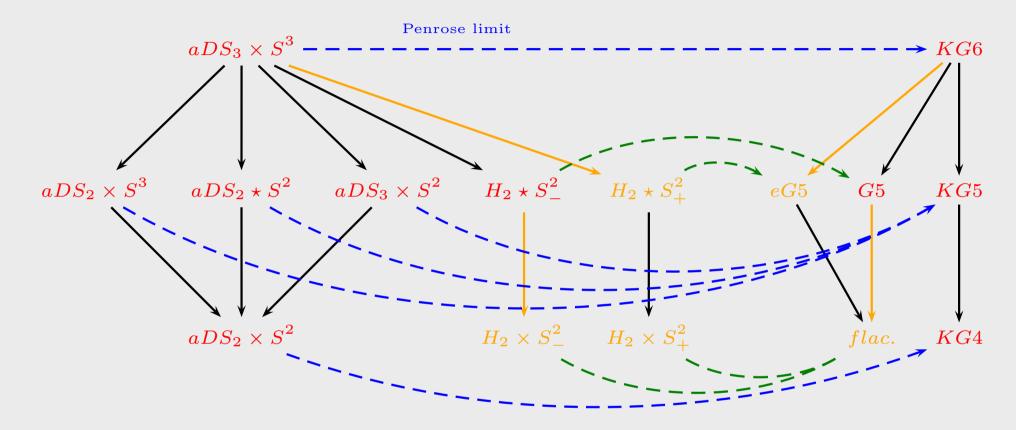
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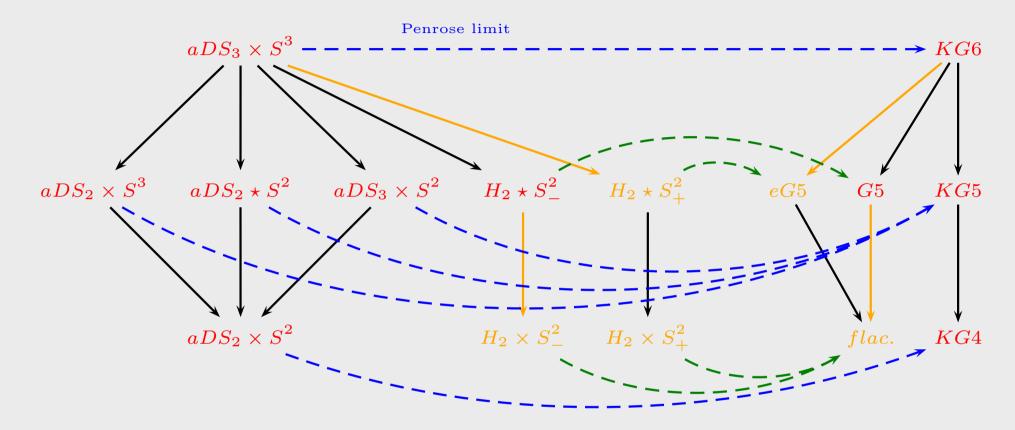
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4 – The Flacuum

As we have seen, the dimensional reduction of the Gödel solution of d = 5, N = 1 SUGRA given by

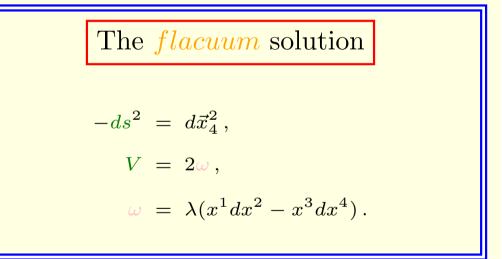
$$(\text{Gödel}) \ G5$$
$$ds^2 = (dt + \omega)^2 - d\vec{x}_4^2,$$
$$V = -\sqrt{3}\omega,$$
$$\omega = \lambda (x^1 dx^2 - x^3 dx^4).$$

IndexSlide 13 / 19
$$\checkmark \bullet$$
SUGRA Vacua1 $\checkmark \bullet$ SUGRA Vacua6 $\checkmark \bullet$ Timelike KK10 \Rightarrow The Flacuum13 \bullet Conclusion18

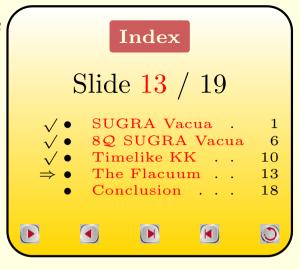




leads to a non-trivial, maximally supersymmetric Euclidean solution of d = 4, N = 2 SUGRA (*i.e.* of the Einstein-Maxwell theory) with flat space and constant **anti-selfdual** field strength *F = -F ($F_{12} = -F_{34} = \lambda/2$)







A constant, anti-selfdual U(1) field strength certainly solves the Maxwell equation in flat space time, but,

how can flat space be a solution in presence of non-trivial matter?

The positivity properties of the action and the energy are opposite in Lorentzian and Euclidean signatures:

	Lorentzian	Euclidean
Action:	$-F^2 = \frac{E^2}{B^2} - \frac{B^2}{B^2}$	$-F^2 = \mathbf{E}^2 + \mathbf{B}^2 > 0$
$T_{\mu u}$:	$F_{\mu}{}^{\rho}F_{\nu\rho} + {}^{\star}F_{\mu}{}^{\rho\star}F_{\nu\rho} > 0$	$F_{\mu}{}^{\rho}F_{\nu\rho} - {}^{\star}F_{\mu}{}^{\rho}{}^{\star}F_{\nu\rho}$

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$$\underline{\text{Lorentzian}} \qquad \underline{\text{Euclidean}}$$

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$$\underline{T_{\mu\nu}:} \qquad F_{\mu}{}^{\rho}F_{\nu\rho} + {}^{\star}F_{\mu}{}^{\rho\star}F_{\nu\rho} > 0 \qquad F_{\mu}{}^{\rho}F_{\nu\rho} - {}^{\star}F_{\mu}{}^{\rho\star}F_{\nu\rho}$$

In particular, selfdual and anti-selfdual Maxwell fields (that can only be defined in Euclidean signature) have a vanishing "energy-momentum" tensor. In general, (anti-) selfdual (non-) Abelian Yang-Mills configurations have vanishing energy-momentum tensors and <u>almost</u> decouple from the metric.

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The decoupling is not complete because (anti-) selfduality $F_{\rho\sigma} = \pm {}^{\star}F_{\rho\sigma}$ has to be proven w.r.t. to a given metric:

$$F_{\rho\sigma} = \pm \frac{1}{2\sqrt{|g|}} g_{\rho\mu} g_{\sigma\nu} \epsilon^{\mu\nu\alpha\beta} F_{\alpha\beta} \,.$$

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 $F = \pm^* F$ and $R_{\mu\nu} = \Lambda g_{\mu\nu}$, then $G_{\mu\nu} + \Lambda g_{\mu\nu} = \frac{1}{2} T_{\mu\nu}$, and $\nabla_{\mu} F^{\mu\nu} = 0$

Page 14-c

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The BPST SU(2) instanton

 $F = \pm^* F$ with any conformally flat metric. Since $F \to 0$ at ∞ we can take that of the round S^4

$$ds^2 = -\frac{d\vec{x}_4^2}{(1+(r/2R)^2)^2}, \quad \Rightarrow \quad R_{\mu\nu} = \frac{1}{R^2}g_{\mu\nu}.$$

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The flacuum U(1) solution

 $F = \pm^* F$ with any conformally flat metric. However, since F is constant, we have to stay with \mathbb{R}^4 which, at most, we can compactify on a torus to have a finite action. $R_{\mu\nu} = 0$ and the Einstein equation is satisfied with zero cosmological constant. Observe that taking the gauge group as U(1) is equivalent to take the time periodic in th eGödel solution.

The vector field of our solution (in a new gauge)

$$V = \lambda (x^{1} dx^{2} - x^{2} dx^{1} - x^{3} dx^{4} + x^{4} dx^{3}) \equiv F_{ab} x^{a} dx^{b},$$

is not strictly periodic on T^4 : when we move around the *a*-th period from x to $x + \hat{a}$ it changes by a gauge transformation

$$V(x+\hat{a}) = V(x) + d\Lambda_a(x), \qquad \Lambda_a(x) = l^{(a)} F_{(a)b} x^b,$$

where $\Lambda_a(x)$ are the U(1) parameters, defined modulo 2π .

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for two integers n, m that label the possible non-trivial bundles. The Euclidean action of the SUGRA solutions is

$$S = -4\pi^2 |\mathbf{nm}| \,.$$

The symmetry superalgebra of the flacuum solution is particularly interesting because it is a deformation of the supertranslation algebra that preserves the commutativity of momenta but modifies slightly the anticommutator of the supercharges (Berkovits and Seiberg)

$$\left\{ \begin{array}{lll} \mathcal{Q}_{(\alpha)}^{\dagger}, \mathcal{Q}_{(\beta)} \right\} &= (\gamma^{1} \gamma^{a})_{\alpha \beta} P_{(a)} & - [\gamma^{1} \frac{1}{2} (1 - \gamma_{5})]_{\alpha \beta} M , \\ \\ \left[\mathcal{Q}_{(\alpha)}, P_{(a)} \right] &= -\mathcal{Q}_{(\beta)} \Gamma_{s} (P_{(a)})^{\beta}{}_{\alpha} , \\ \\ \left[\mathcal{Q}_{(\alpha)}, M \right] &= -\mathcal{Q}_{(\beta)} \Gamma_{s} (M)^{\beta}{}_{\alpha} , \\ \\ \left[P_{(a)}, M \right] &= -P_{(b)} \Gamma_{v} (M)^{b}{}_{a} , \end{array}$$

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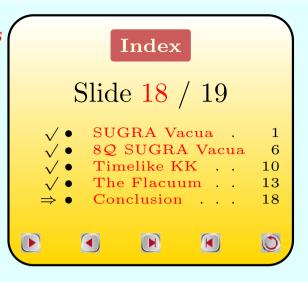
This superalgebra can be obtained by dimensional reduction of the Gödel superalgebra, in which the momenta $P_{(a)}$ do not commute, but give $P_{(0)}$ which should be interpreted as the generator of U(1) gauge transformations on d = 4. This property is, precisely, what allowed us to relate the periods of the torii.

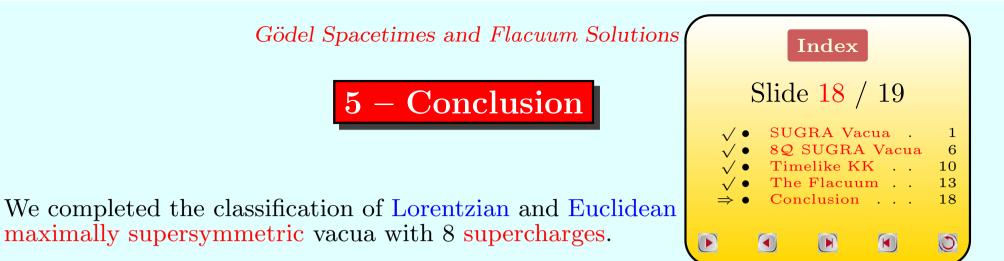




5 – Conclusion

 \star We completed the classification of Lorentzian and Euclidean maximally supersymmetric vacua with 8 supercharges.





 \star We have found a solution, the *flacuum* solution with very interesting properties and that can be generalized to other dimensions (always as a timelike reduction of a Gödel-type solution).

Gödel Spacetimes and Flacuum Solutions 5 - Conclusion Index \$\screwtyle{2}\$ - Conclusion \$\screwtyle{2}\$ Slide 18 / 19 \$\screwtyle{2}\$ - SUGRA Vacua \$\screwtyle{2}\$ SUGRA Vacua \$\screwtyle{2}\$ - SUGRA Vacua \$\screwtyle{2}\$ SUGRA Vacua \$\screwtyle{2}\$ - Sugra Vacua \$\screwtyle{2}\$ SUGRA Vacua \$\screwtyle{2}\$ - Timelike KK \$\screwtyle{2}\$ - The Flacuum ...

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- ★ We have not discussed how the compactification affects the residual supersymmetry of the solution, which is a delicate point because the holonomy of the solution is not contained in SO(4).
- ★ We have determined the symmetry superalgebra of the *flacuum* solution. We notice that the symmetry superalgebras of all the maximally supersymmetric vacua are always deformations of the supertranslation (superPoincaré) algebra, which may allow to classify and find all these vacua.

